



# Energy release processes during the 17 May 2013 M3.2 solar flare



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3 -Moscow MV Lomonosov State University, Russia

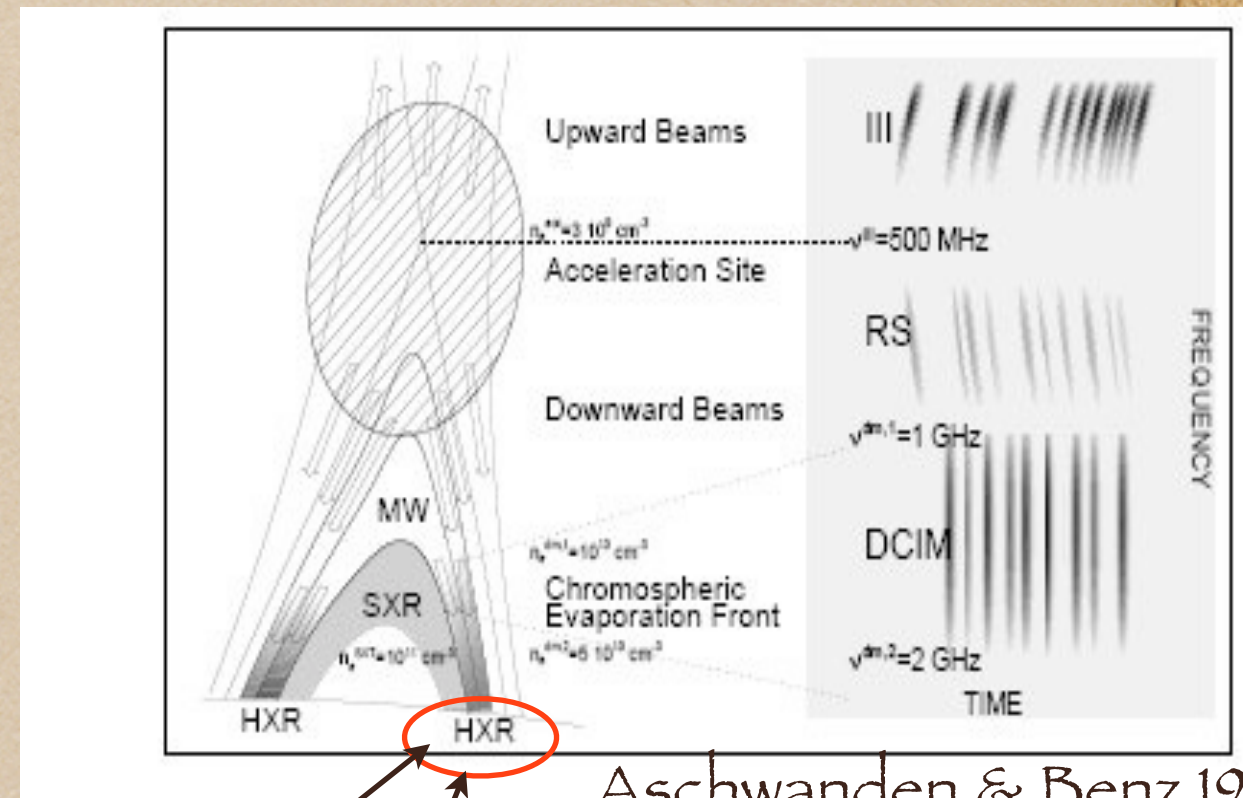
4-St.Petersburg Branch of SAO RAS, St. Petersburg, Russia



# Motivations

- ◆ Flares with several peaks observed in HXR and MW temporal profiles What is it? Series of reconnection or development of single energy input process modulated by some oscillations?

Why to use H- $\alpha$  data?



Aschwanden & Benz 1977

UV continuum

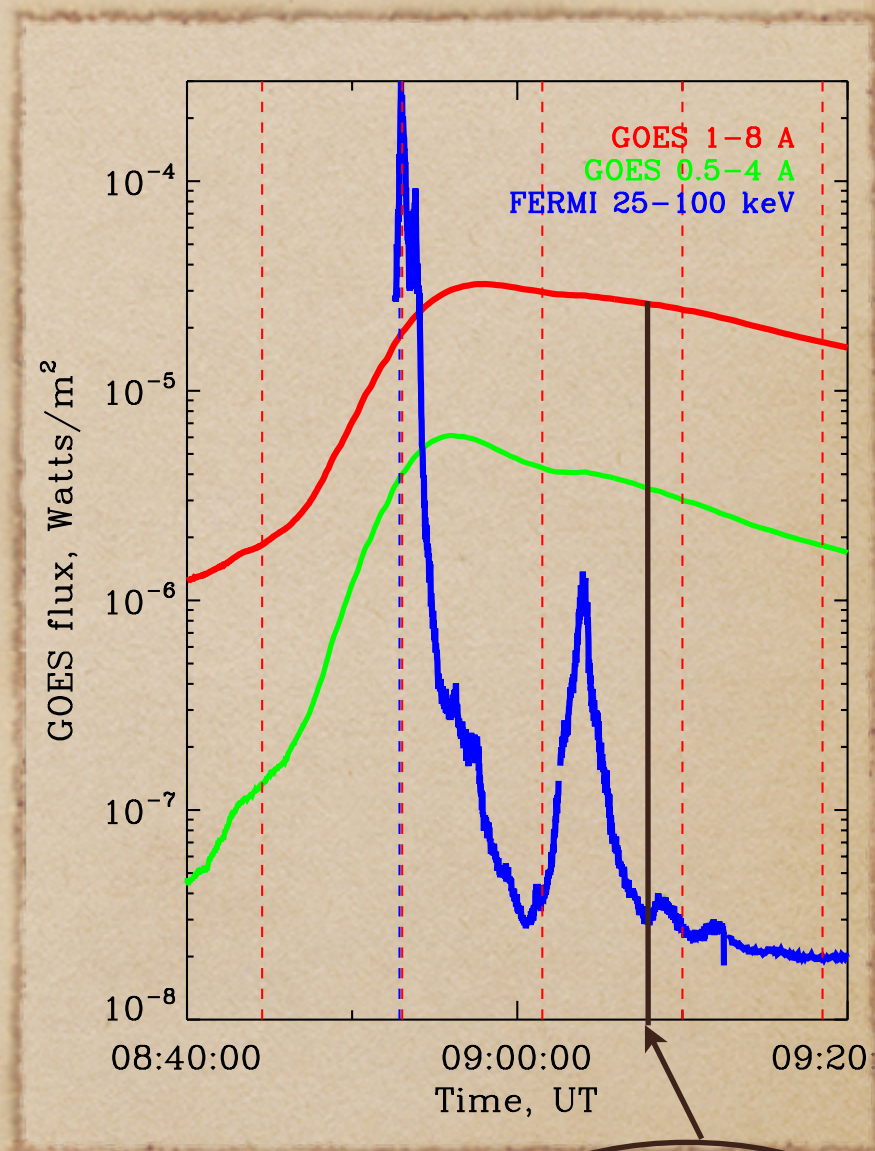
H- $\alpha$  emission



# Event - M3.2 flare

## SOL2013-05-17T08:43

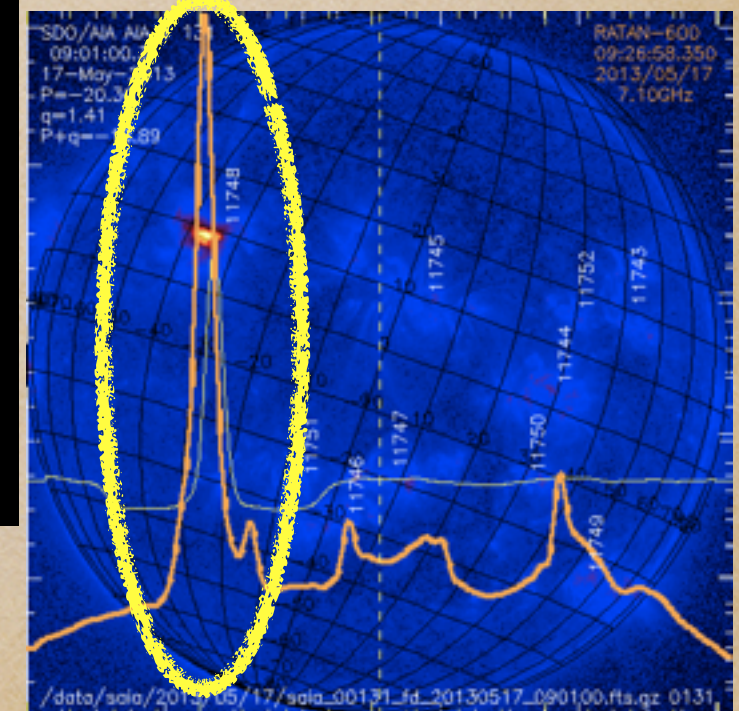
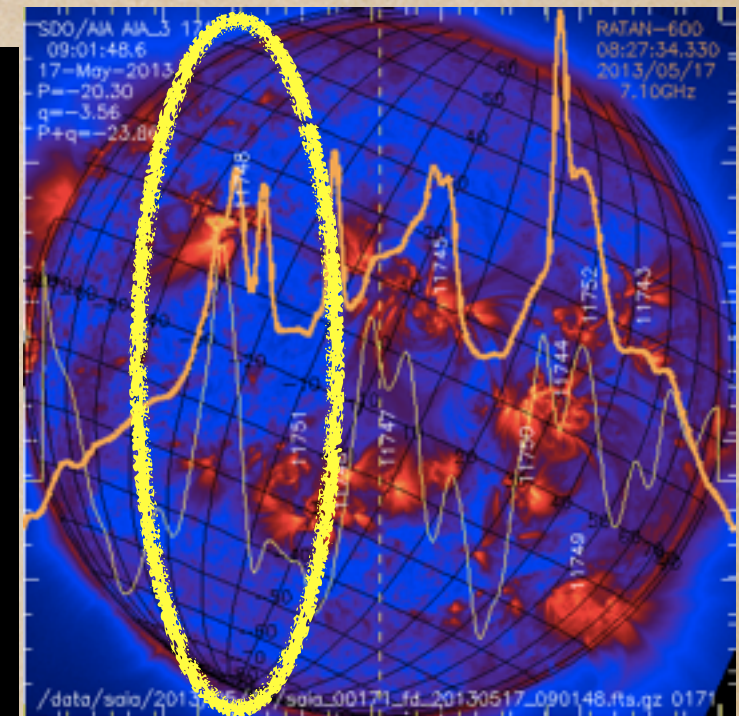
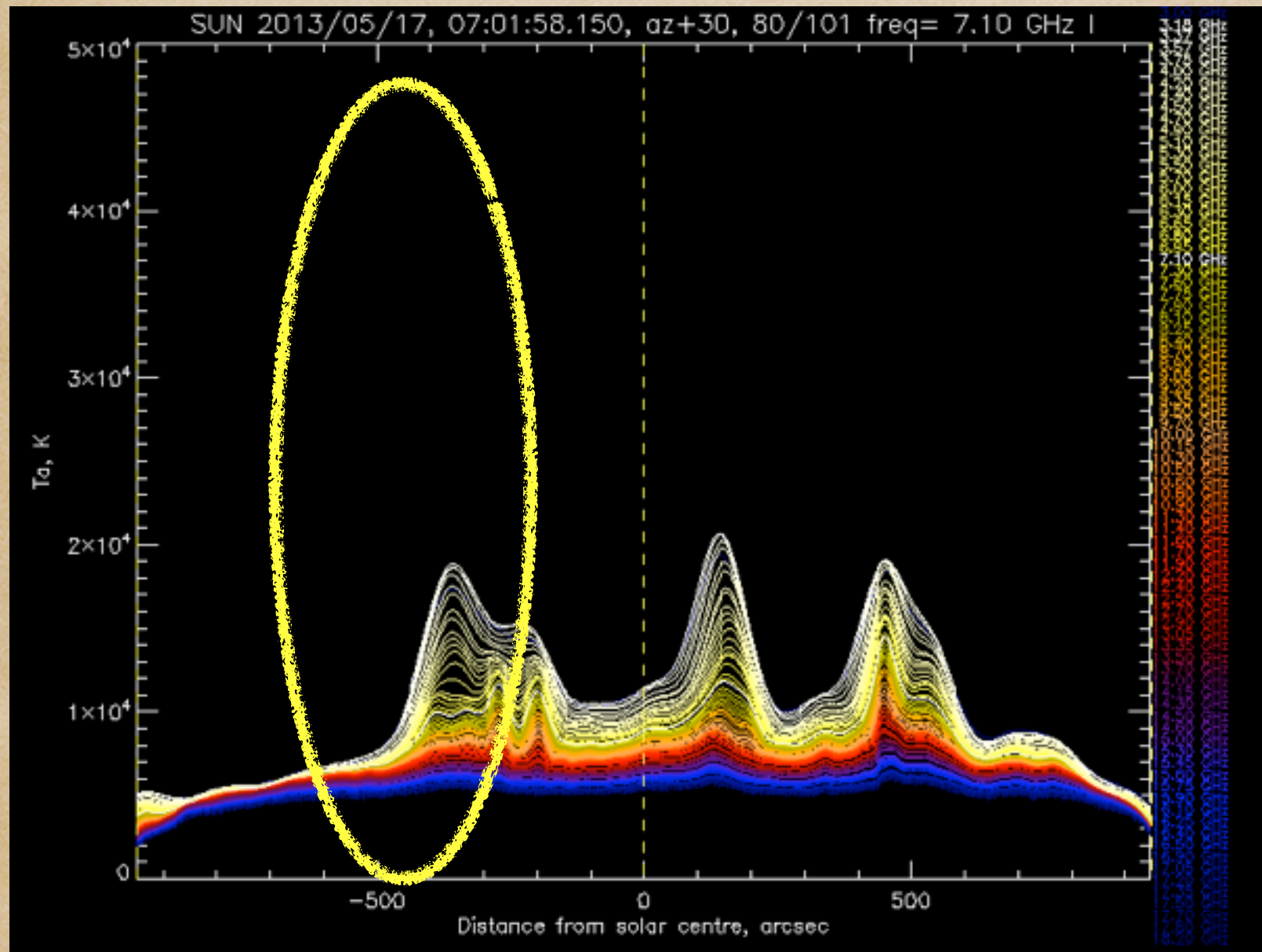
- ◆ HSFA-2 (Ondřejov):  $H_{\alpha}$  line
- ◆ RATAN-600 : MW data
- ◆ The Siberian Solar Radio Spectropolarimeters (2-24 GHz & 4-8GHz): MW spectra
- ◆ RHESSI & FERMI : X-ray data
- ◆ AIA/SDO and HMI/SDO : EUV and magnetograms



RHESSI  
observations



# RATAN-600 data

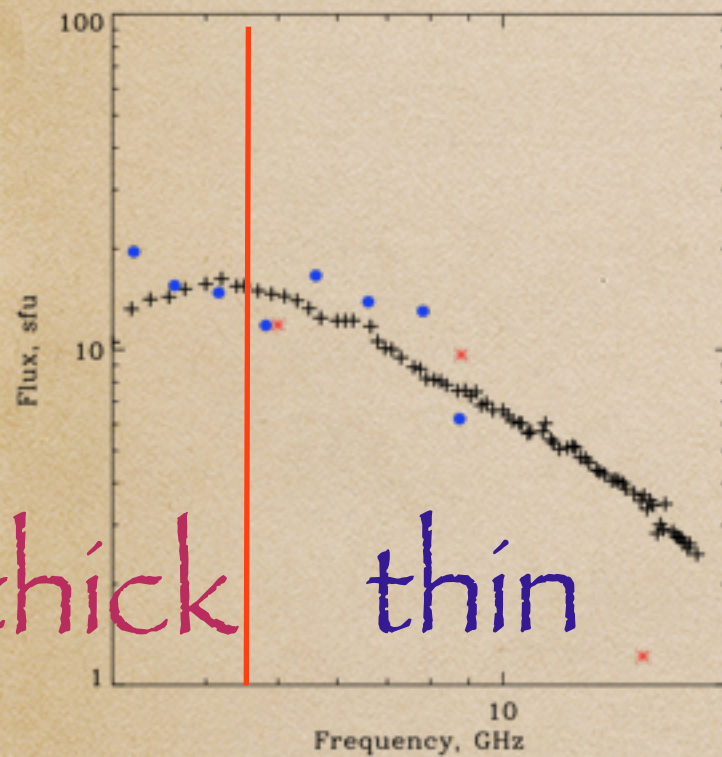




# MW spectra

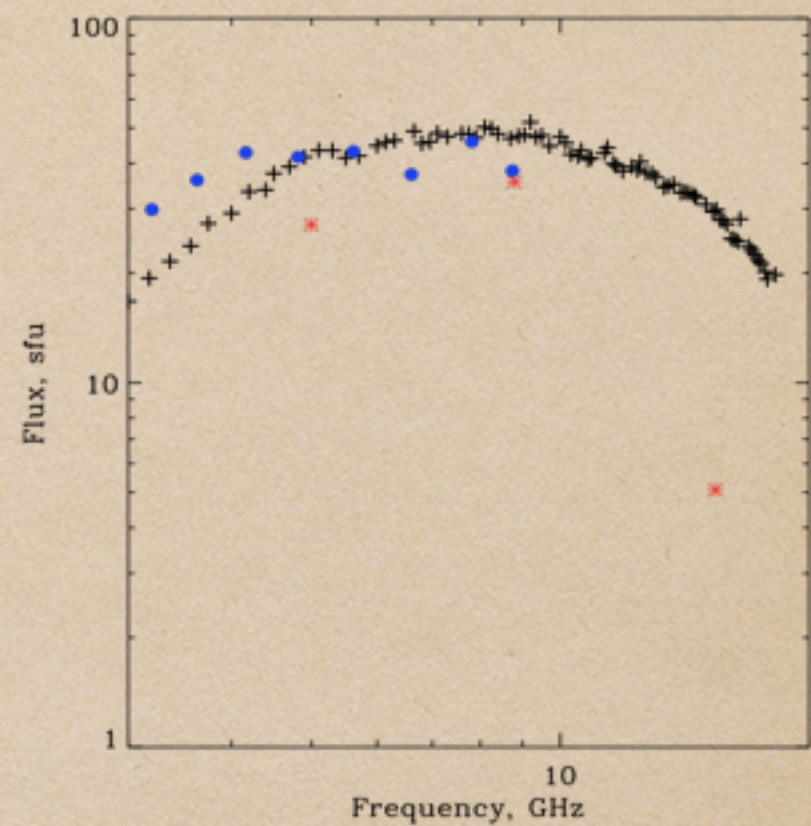
08:44 UT

09:19 UT

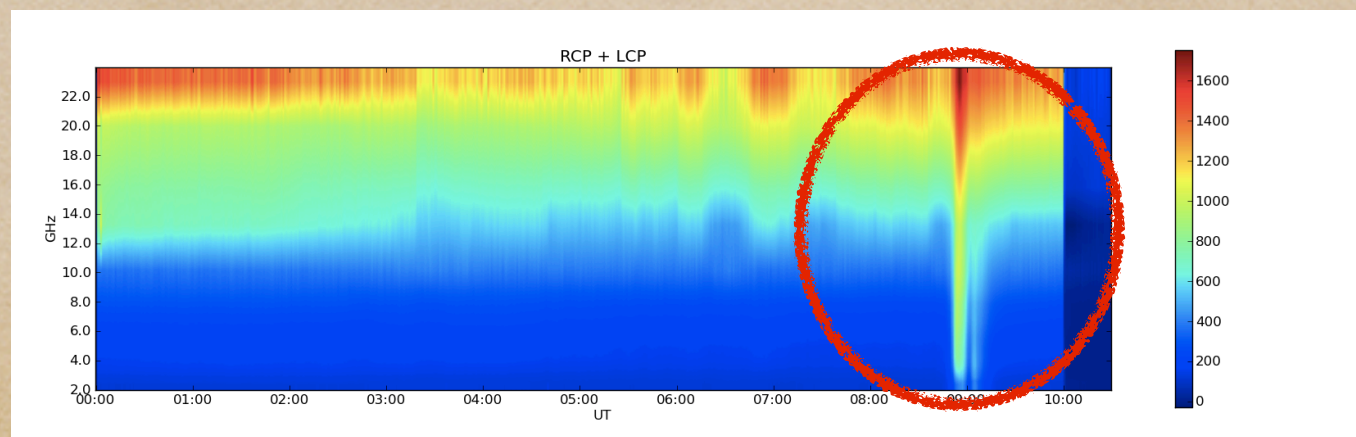


thick thin

RATAN-600  
SSRS  
RSTN



SSRS 2-24 GHz

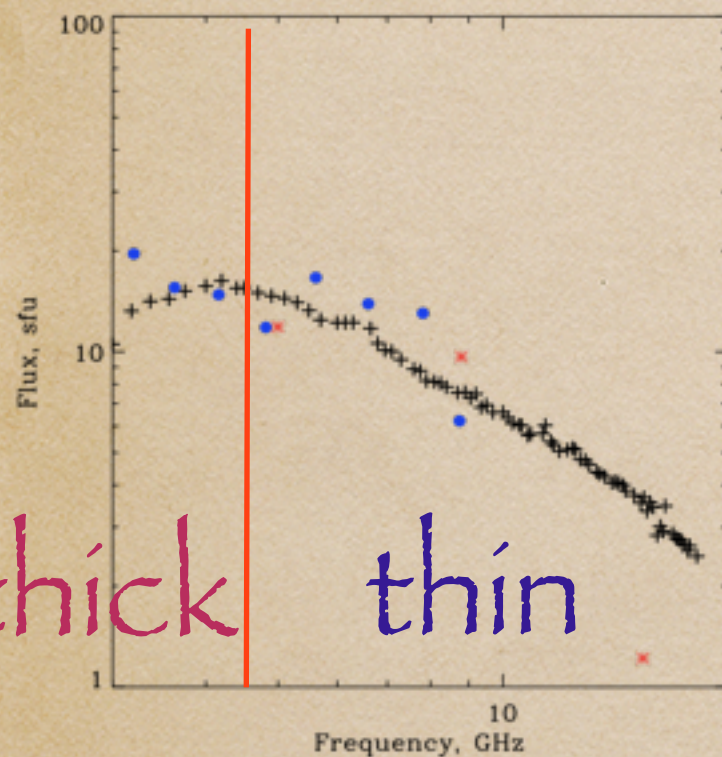




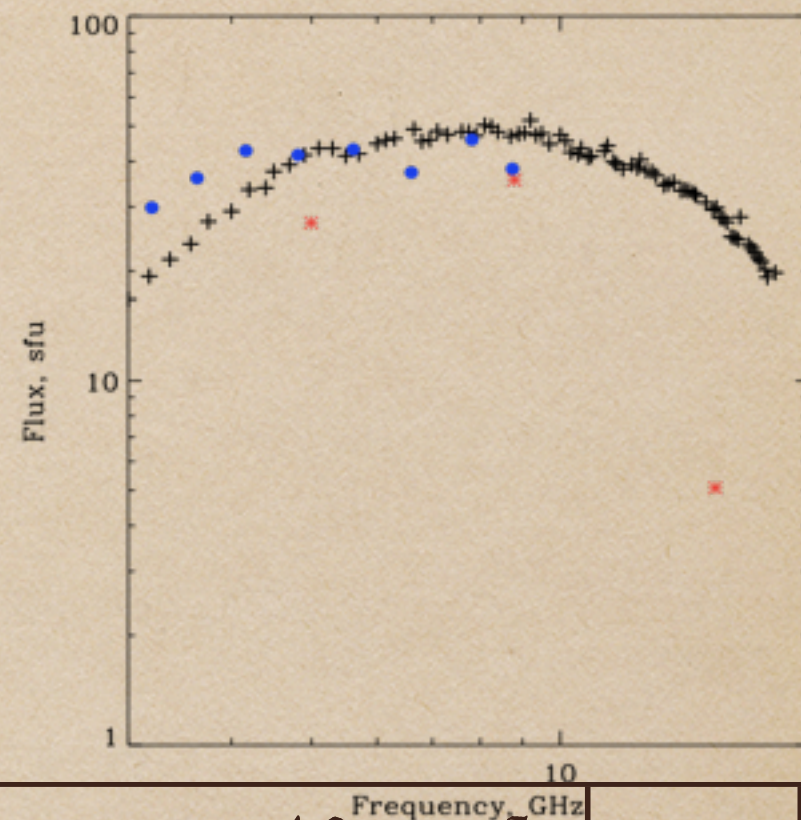
# MW spectra

08:44 UT

09:19 UT



RATAN-600  
SSRS  
RSTN



Time, UT	Te, MK	ME, $10^{49} \text{ cm}^{-3}$	$\gamma$
08:53	21	0.14	3.3
09:04	21	0.09	3.0
9:08	20	0.08	4.1
09:12	19	0.07	~



# H $\alpha$ -image processing method

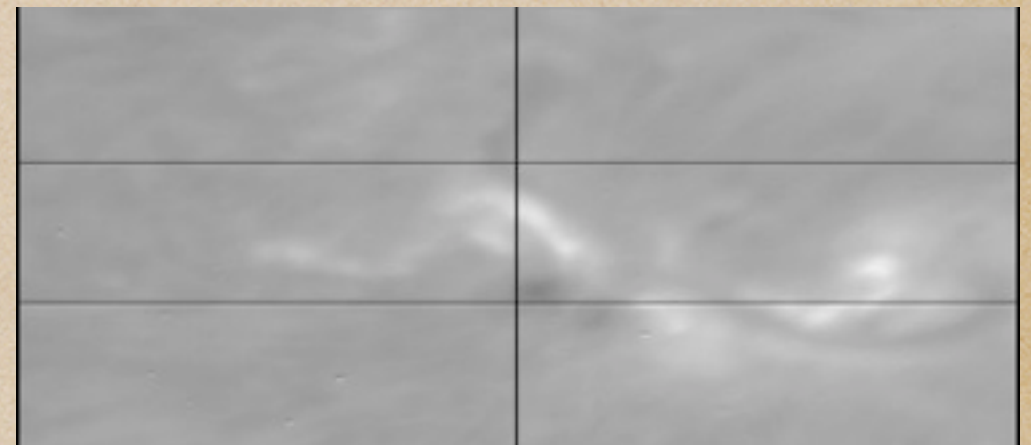
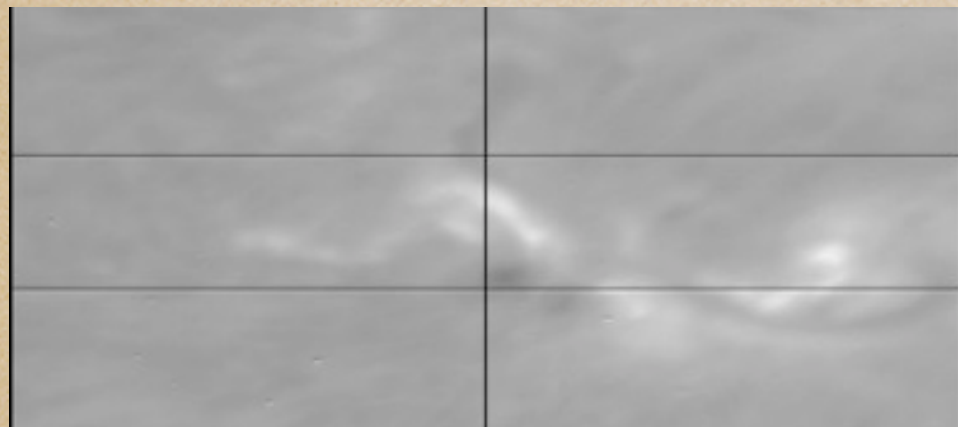
H $\alpha$  slit-jaw images (halfwidth  $\sim 0.5 \text{ \AA}$ )

Spatial evolution

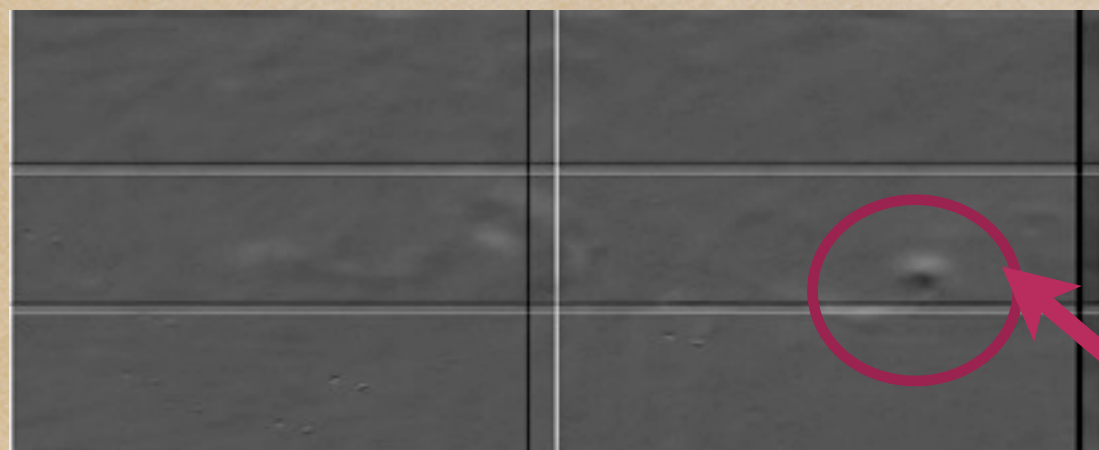
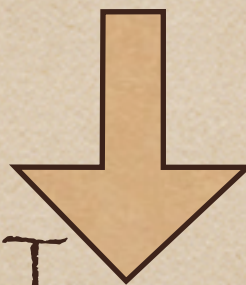
H $\alpha$  intensity evolution

08:43:37 UT

08:43:36 UT



08:43:37 - 08:43:36 UT



Kernel 1

The method was tested  
in Zharkova et al  
MNRAS, 2011



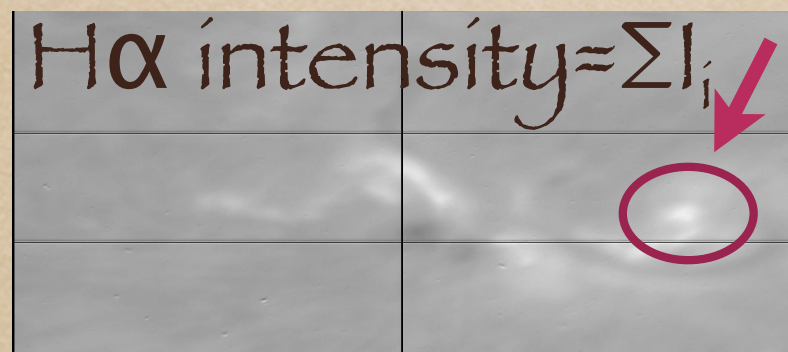
# H $\alpha$ -image processing method

H $\alpha$  slit-jaw images (halfwidth  $\sim 0.5 \text{ \AA}$ )

Spatial evolution

H $\alpha$  intensity evolution

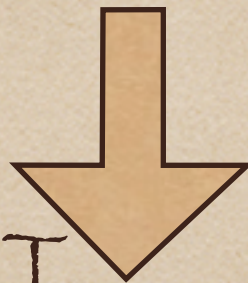
08:43:37 UT



08:43:36 UT



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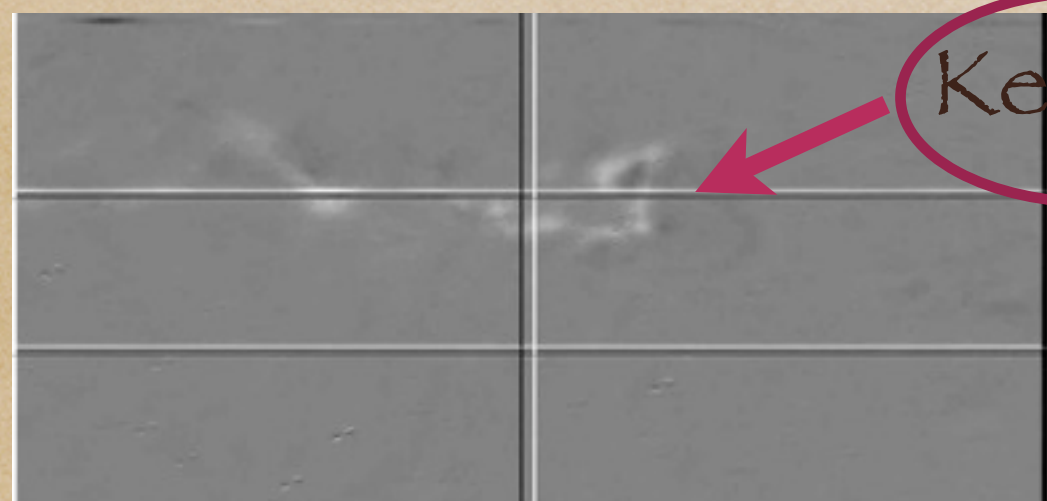
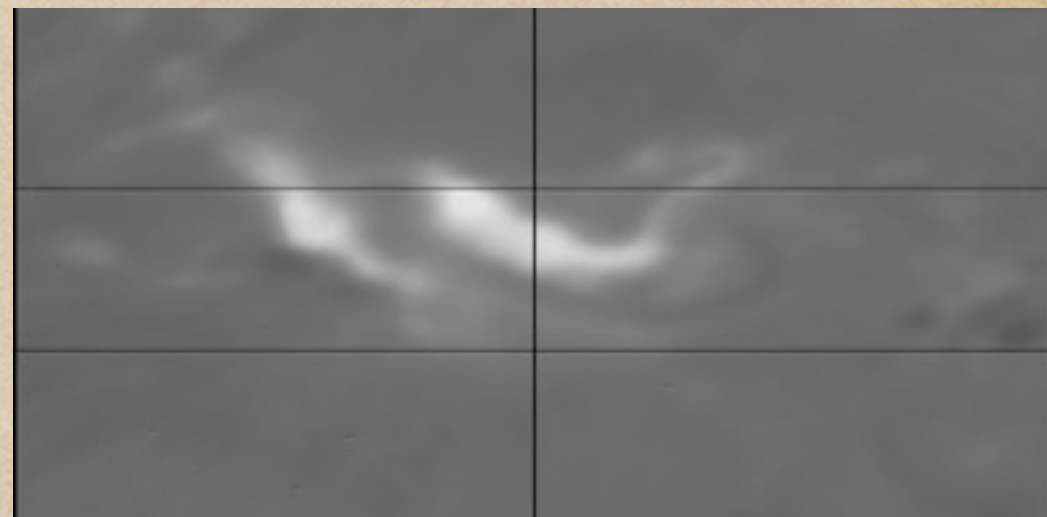


# Temporal evolution of $H\alpha$ emission

08:47:58 UT



08:53:35 UT

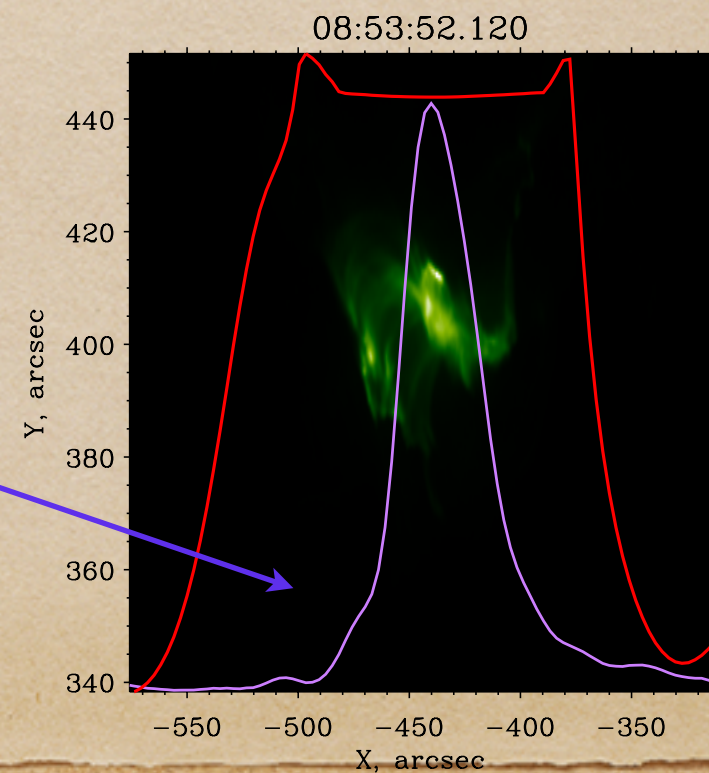
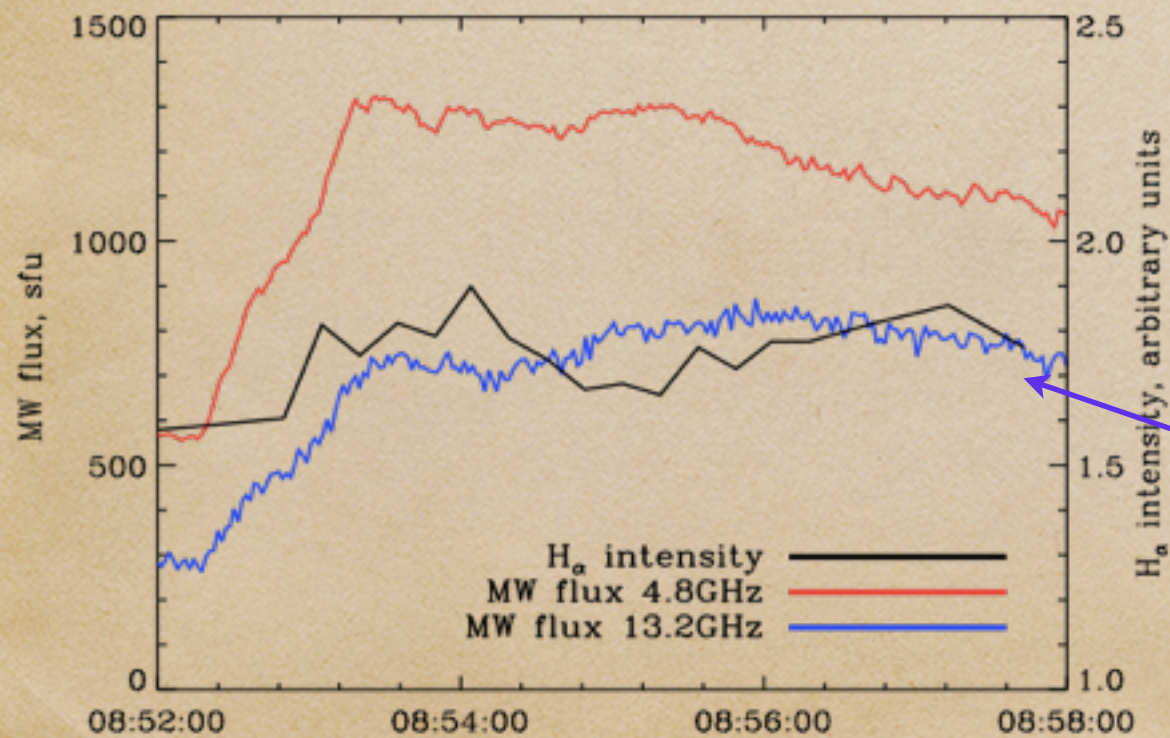
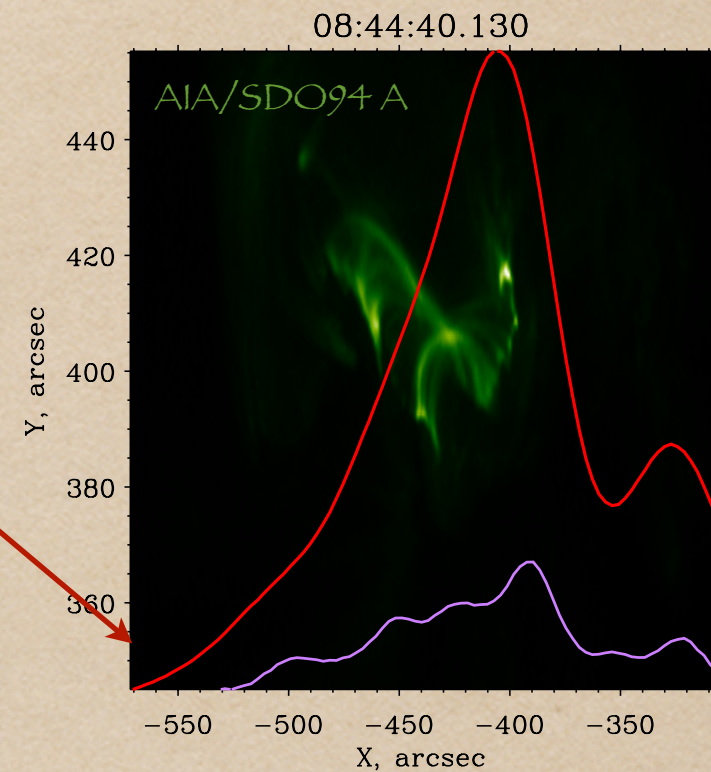
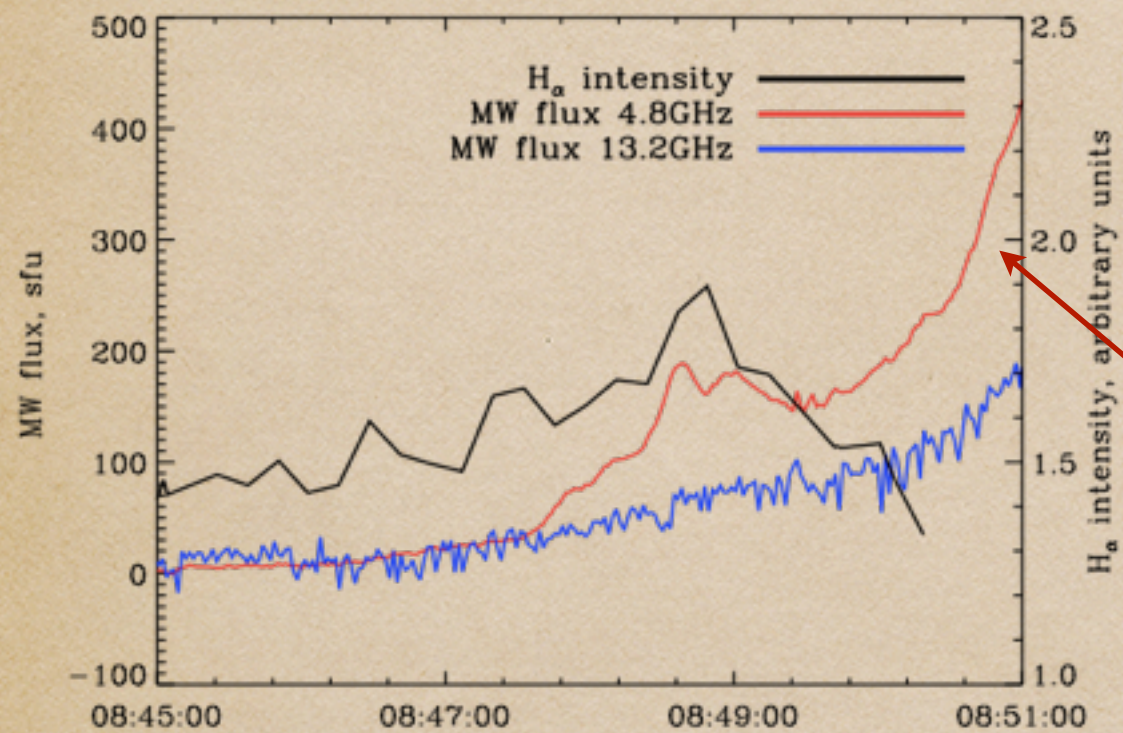


Kernels



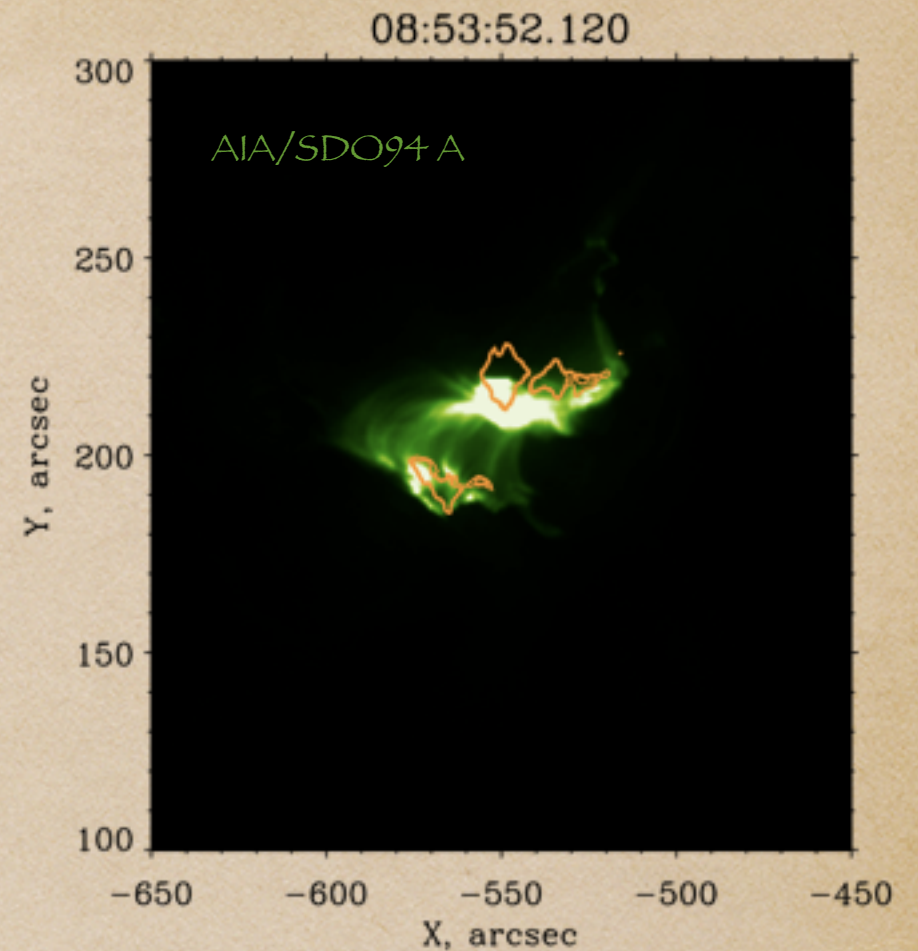
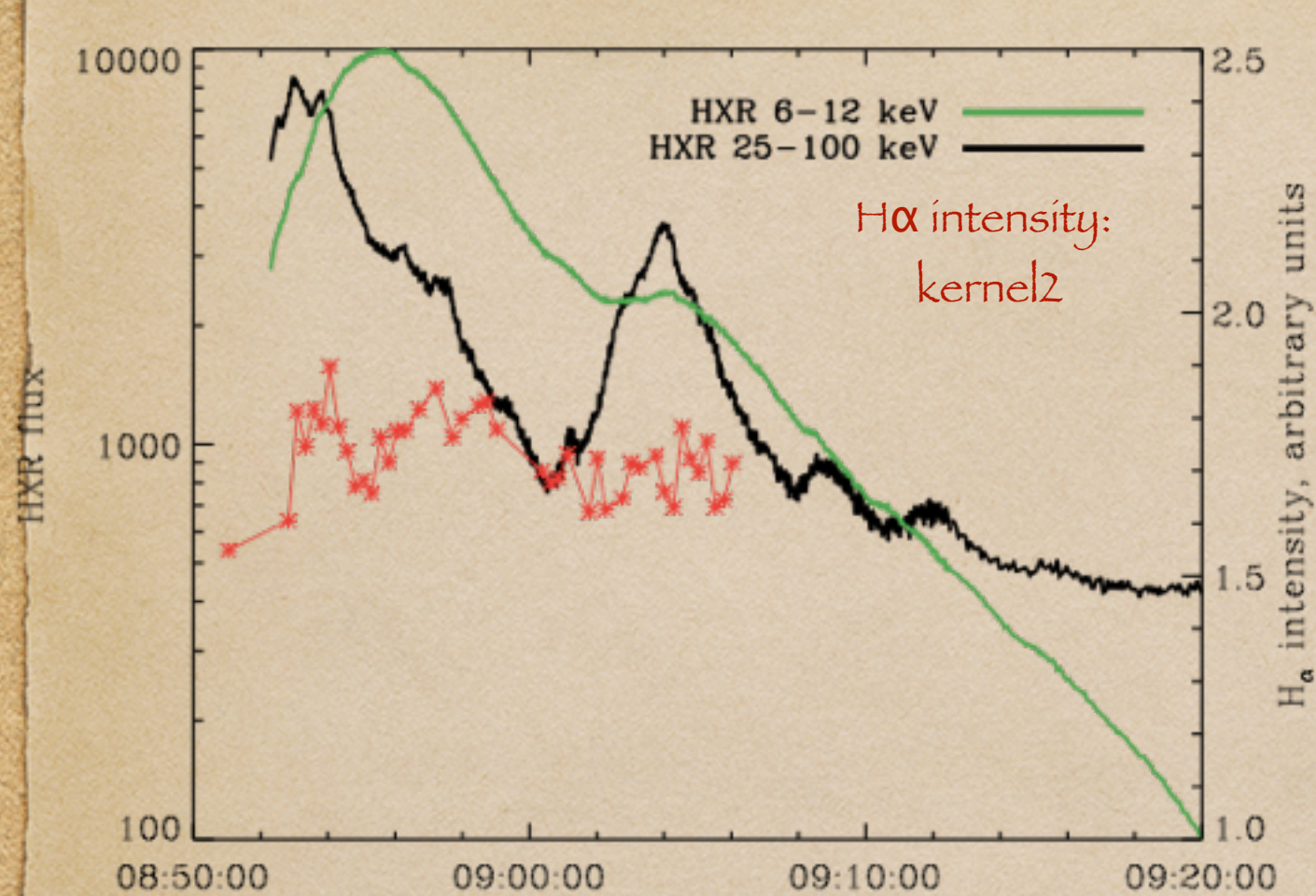


# H $\alpha$ vs MW: thermal or not-thermal?





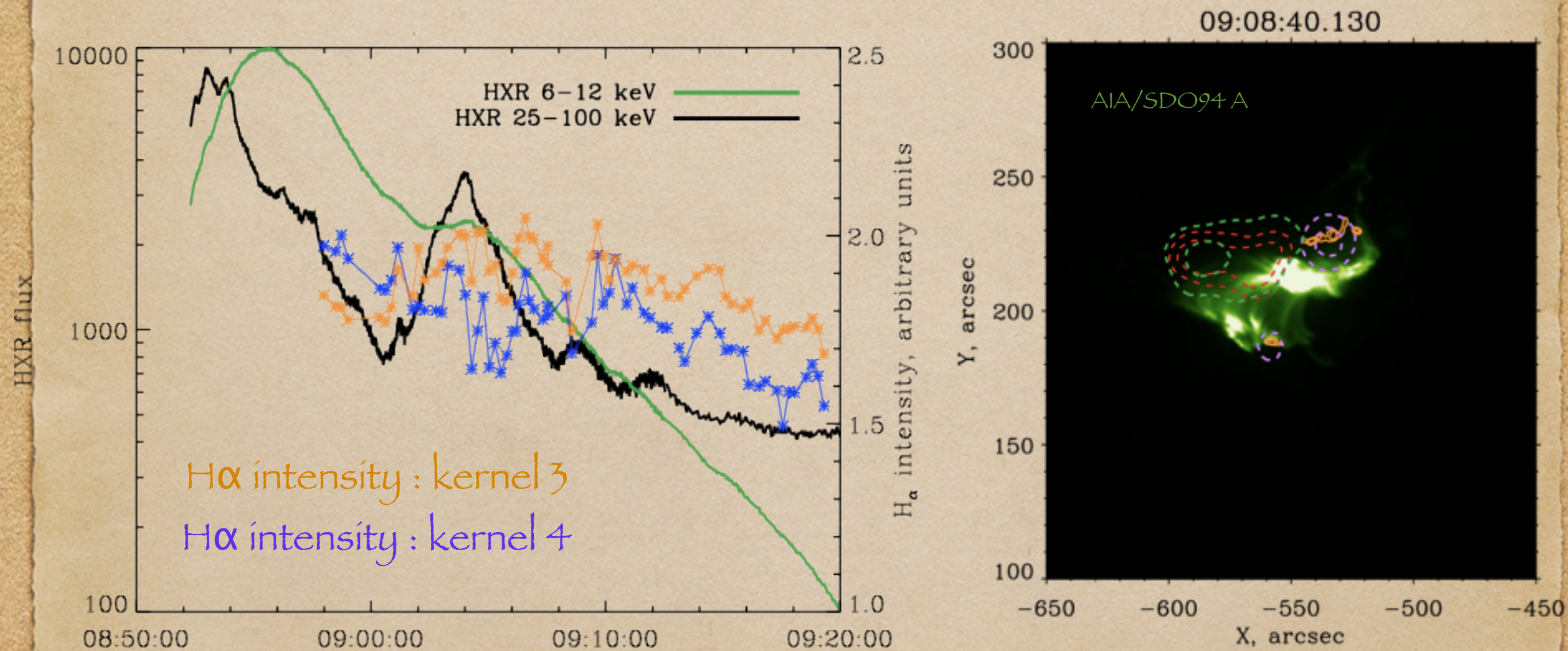
# H $\alpha$ kernels vs HXR



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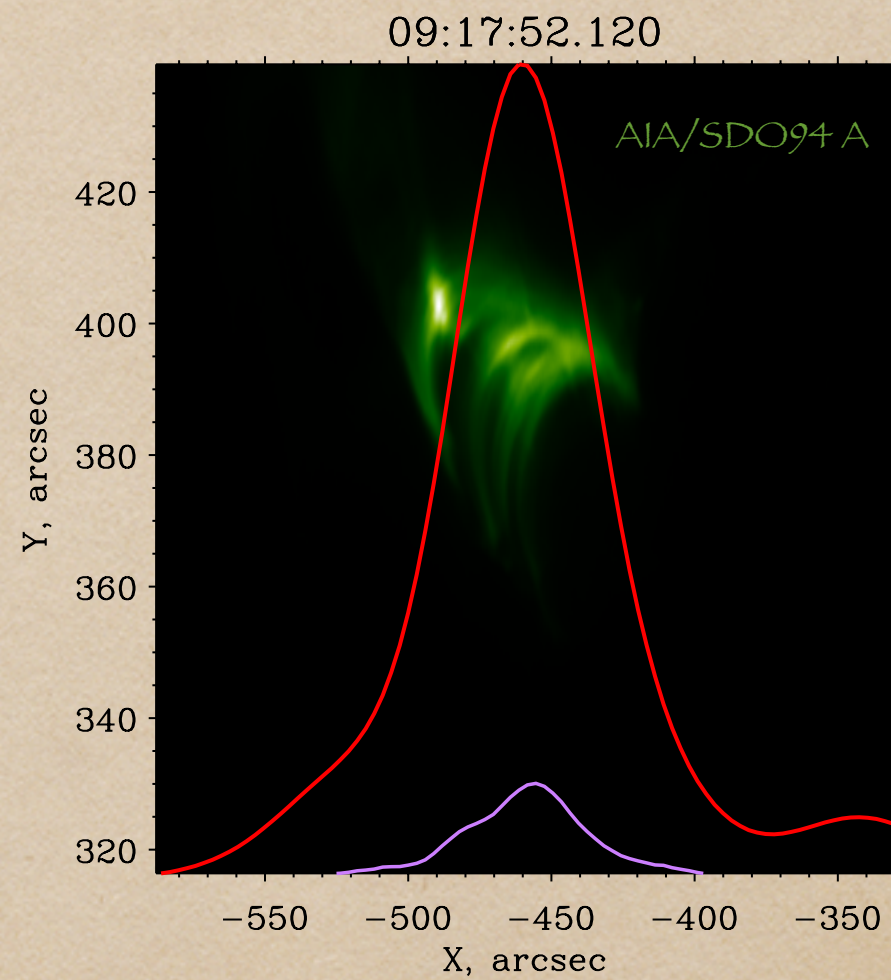
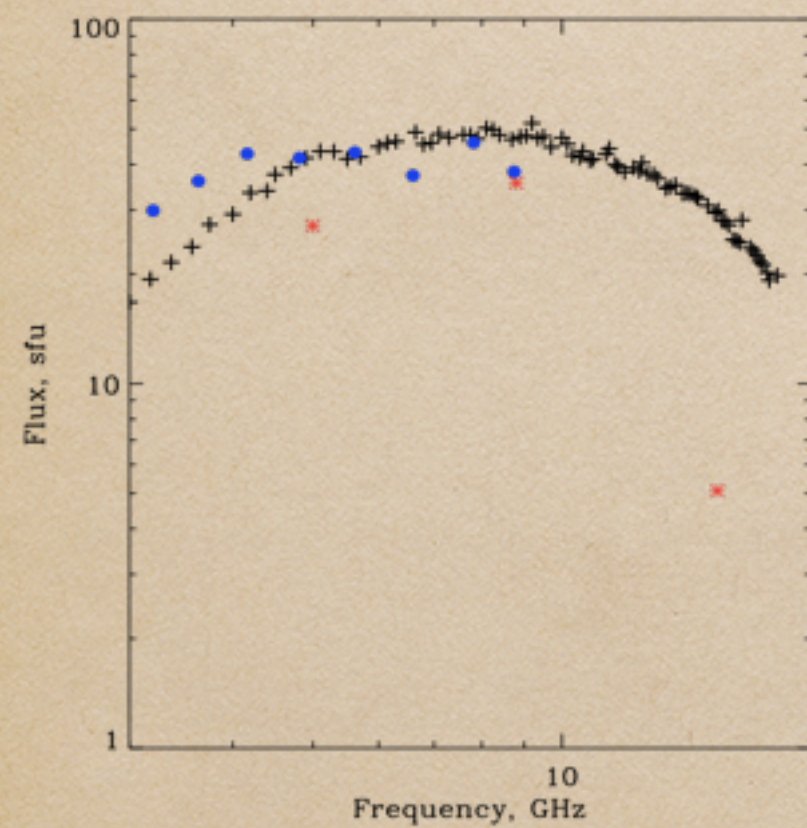


# H $\alpha$ kernels vs HXR





09:19 UT





# Summary

- ◆ It was revealed that the individual flare H $\alpha$  kernels that arose in different moments of the flare were dominated by different mechanisms.
- ◆ Mainly they arose as a result of accelerated electrons precipitation into lower layers of solar atmosphere. Further development these kernels could be described by different thermal processes.



# Summary II

- ◆ The MW spectra observed after maximum of the flare was a result of superposition of at least two MW sources or flare loops.



Thank you for attention!