Search for signature of the LIGO gravitational wave events in SPI-ACS and GBM/Fermi

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Outline

- No evident GRB-like event is found for published LIGO GW150914, GW151226, LVT151012, GW170104
- Most sensitive omnidirectional GRB experiments = GBM/Fermi and SPI-ACS/INTEGRAL
- Searching for quasi-periodic transient events
- Special case of GW150914
 - comparison of sensitivity for short GRBs
- How to reconcile GBM/Fermi and SPI-ACS observations?
- BH-BH merging: what could be a model EM-emission?

Searching for quasi-periodic transient events

- As a part of time series investigation we developed a code for continuous search QPO-like events localized in time.
- We apply the code to SPI-ACS ...

SPI-ACS/INTEGRAL GW 150914 wavelet (Morlet) intermittency scalogram



GW150914

Power Density Spectrum P = 7.1 s



v, Hz

SPI-ACS time profile P=7.1 s



SPI-ACS/INTEGRAL GW 150914-QPO wavelet (Morlet) intermittency scalogram



GW150914

Chance probability?

1x10⁻⁵ – 2x10⁻⁴, depending on the estimating procedure

GBM/Fermi scalogram?

 No significant feature analogous to GW 150914-QPO is detected in GRB/Fermi data at time of GW 150914-GBM

GRB/Fermi GW 150914-GBM (Connaughton+, 2016)

GBM detectors at 150914 09:50:45.797 +1.024s



GW 150914-GBM

- A typical short GRB?
- Should be detected by SPI-ACS?
- GW 150914-GBM may be detected by SPI-ACS at ~5σ based on our statistical comparison (cf. 15σ deduced by Savchenko+, 2016)

Statistical SPI-ACS and GBM/Fermi sensitivity comparison

- SPI-ACS short GRBs sample
- Short GRBs of GBM catalog
- Log*N* Log*S* normalization assuming parent population is the same for both experiment

Statistical SPI-ACS and GBM/Fermi sensitivity comparison



Statistical SPI-ACS and GBM/Fermi sensitivity comparison



SPI-ACS/INTEGRAL sky map

Localization of GW 150914-GBM

(Connaughton+, 2016)





How to evenly reconcile GBM/Fermi and SPI-ACS observations? (0)

 Let suggest that GW 150914-GBM and GW 150914-ACS/QPO are real astrophysical events How to evenly reconcile GBM/Fermi and SPI-ACS observations? (I)

- GW 150914-GBM and GW 150914-ACS/QPO have a different nature
- GW 150914-GBM is a usual short GRB occulted by the Earth for SPI-ACS FOV and not related to GW150914
- Actual position of GW150914 EM counterpart (which registered as 150914-ACS/QPO) is occulted by the Earth for GBM/Fermi
- Complicated? But not impossible

How to evenly reconcile GBM/Fermi and SPI-ACS observations? (II)

- GW 150914-GBM and GW 150914-ACS/QPO have the same nature
- GW 150914-GBM is a short GRB related to GW150914
- GW 150914-ACS/QPO is an extended emission sometime observed after short GRBs
- For some reasons GBM/Fermi not detected GW150914-ACS/QPO and SPI-ACS not detected GW150914-GBM
- Actual position of GW150914 EM counterpart is significantly restricted by small part of the LIGO localization error box

Thank you

